

Time-dependent BPS configurations in String and M-theory

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“Dark Universe”, Su-Zhou 2006-11-2

Based on hep-th/0508191, 0509113 (with Ya-li He, Peng Zhang) and 0603147 (with Hong-zhi Chen).

1 Motivations

- Singularity theorem: (Penrose, Hawking ... 1960s) Singularities could be ill defined in General Relativity. Especially, spacelike singularities are BAD: Black holes, big-bang (crunch) singularities;
- Very near singularities: quantum effects should be taken into account. The concept of geometry must be broken down. So what? Quantum foam, spin foam, It's a question for every kind of quantum gravity theory!

- The resolution of cosmic singularity? One of the central issues in String theory. Singularity should be resolved in string theory due to the extensive nature of string.
 - Orbifold singularity: resolved by perturbative D.O.F.
Dixon et.al., ...
 - Conifold singularity: resolved by nonperturbative D.O.F.
A. Strominger

- What are the right degrees of freedom?
 - Perturbative description? Null orbifolds (suffer from the large blueshift) [H. Liu et.al.](#), [Horowitz et.al.](#),... Milne universe [Berkooz et.al.](#), [Nekrasov](#), ...
 - Closed string Tachyon phase? (poorly understood) [E. Silverstein](#), ...
 - Pre-big-bang scenario? [Veneziano et.al.](#)
 - Nonperturbative description? Matrix cosmology!

- **Matrix cosmology:** C. Craps, Sethi and Verlinde (2005)
 - IIA flat spacetime with a linear null dilaton background;
 - 1/2 BPS;
 - Perturbatively solvable;
 - Geodesics incomplete;
 - String coupling diverges near the big-bang singularity;
 - A dual Matrix string description (2-dim SYM on Milne orbifold);
 - 1-loop effective potential? Miao Li and Wei Song, Craps et.al.

- Questions: other time-dependent configurations with SUSY
 - 1/2 BPS configurations in M-theory?
Miao Li, Bin Chen, N. Ohta ...
 - Other solvable BPS background in String theory?
S. Das et.al, B. Chen et.al.

– Configurations with holographic description?

Chu and Ho, S. Das et.al., F.L. Lin et.al.

Holographic description of BH singularity via AdS/CFT

Shenker et.al....

Emergent time?

2 Time-dependent BPS configurations in M-theory and Matrix models

Start from 11-dimensional supergravity

$$R_{MN} = \frac{1}{12}(F_{MPQR}F_N{}^{PQR} - \frac{1}{12}g_{MN}F^2)$$
$$d * F = \frac{1}{2}F \wedge F$$

and the Killing spinor equations

$$\tilde{D}_M \epsilon = (D_M - \Omega_M) \epsilon$$

where D_M is the spin connection defined by

$$D_M = \partial_M + \frac{1}{4}\omega^{ab}\Gamma_{ab}$$

and

$$\Omega_M = \frac{1}{288} F_{PQRS} (\Gamma^{PQRS}_M + 8\Gamma^{PQR} \delta_M^S).$$

We would like to find the solutions to the equations of motions, which have at least sixteen standard supersymmetries, corresponding to the Killing spinors ϵ satisfying $\Gamma^+ \epsilon = 0$.

- **Vacuum configurations:** we make the following ansatz:

$$ds^2 = 2A_0(u)dudv + A_i(u)(dx^i)^2,$$

where the $A_0(u), A_i(u), i = 1, \dots, 9$ are the functions of u .

An orthogonal frame

$$\theta^+ = \sqrt{A_0(u)}du, \quad \theta^- = \sqrt{A_0(u)}dv \quad \theta^I = \sqrt{A_i(u)}dx^i \delta_i^I.$$

The nonvanishing spin connections are

$$\omega^{-i} = -\frac{\partial_u \sqrt{A_i}}{\sqrt{A_0}} dx^i, \quad \omega^{-+} = -\frac{\partial_u \sqrt{A_0}}{\sqrt{A_0}} du$$

and the only non-zero Ricci tensor is

$$R_{uu} = \sum_i \frac{\sqrt{A_0}}{\sqrt{A_i}} \left(-\partial_u \left(\frac{\partial_u \sqrt{A_i}}{\sqrt{A_0}} \right) + \frac{\partial_u \sqrt{A_i} \partial_u \sqrt{A_0}}{A_0} \right).$$

Next let us check the remaining supersymmetries. In the vacuum case, $\Omega_M = 0$ and the Killing spinor equations are

$$\begin{aligned}\partial_u \epsilon &= -\frac{1}{2}(\omega_u^{+-} + \omega_u^{-+} \Gamma_{-+} + \omega_u^{-i} \Gamma_{-i}) \epsilon \\ \partial_i \epsilon &= -\frac{1}{2} \omega_i^{-k} \Gamma_{-k} \epsilon.\end{aligned}$$

Choosing a constant spinor ϵ_0 with $\Gamma^+ \epsilon_0 = 0$, then the Killing spinor satisfying the above equations is

$$\epsilon(u) = \exp\left(-\frac{1}{2} \int \omega_u^{-+} du\right) \epsilon_0.$$

- Plane-wave like configurations

Let us introduce a constant 4-form field strength

$$F_{u123} = f_0(u),$$

and make the following ansatz on the metric

$$ds^2 = 2A_0(u)dudv + B_0(u, x)du^2 + A_i(u)(dx^i)^2 + B_i(u, x)dx^i du,$$

where the B_0, B_i 's are the functions of u and x . **Very recently, N. Ohta et.al. generalized the above metric and found more general supersymmetric configurations.**

The metric allows an orthogonal frame

$$\theta^+ = \sqrt{A_0(u)} du$$

$$\theta^- = \sqrt{A_0(u)} dv + \frac{B_0(u, x)}{2\sqrt{A_0(u)}} du + \frac{B_i(u, x)}{2\sqrt{A_0(u)}} dx^i$$

$$\theta^I = \sqrt{A_i(u)} dx^i \delta_i^I.$$

The corresponding spin connections are

$$\omega^{-+} = -\frac{\partial_u \sqrt{A_0}}{\sqrt{A_0}} du$$

$$\omega^{+i} = 0$$

$$\omega^{ij} = -\frac{\partial_i B_j}{2\sqrt{A_i A_j}} du$$

$$\omega^{-i} = \frac{1}{\sqrt{A_i}} \left(\frac{\partial_i B_0}{2\sqrt{A_0}} - \frac{\partial_u B_i}{2\sqrt{A_0}} + \frac{\partial_u \sqrt{A_0}}{A_0} B_i \right) du$$

$$- \frac{\partial_u \sqrt{A_i}}{\sqrt{A_0}} dx^i + \sum_{j \neq i} \frac{\partial_i B_j}{2\sqrt{A_0 A_i}} dx^j.$$

The existence of Killing spinor with $\Gamma^+ \epsilon = 0$ requires that

$$B_i = A_{ij}(u)x^j$$

with $A_{ij} = -A_{ji}$ ¹. And the equation of motion asks B_0 to be bilinear in x^i , namely

$$B_0 = B_{ij}(u)x^i x^j.$$

Therefore, we find a class of time-dependent configurations akin to the plane-waves

$$ds^2 = 2A_0(u)dudv + B_{ij}(u)x^i x^j (du)^2 + A_i(u)(dx^i)^2 + A_{ij}(u)x^j dx^i du$$

with at least 16 supersymmetries.

¹The antisymmetric condition is not necessary

- Extra supersymmetries?
 - Why? The "standard" 16 supersymmetries lead to nonlinearly realized supersymmetries on Matrix model (or on the action of other objects embedded in the background).
 - Strategy: check if such configurations have Killing vector with nonvanishing ∂_u component and no dependence on v .

1. Consider the Killing vector

$$K = \bar{\epsilon} \Gamma^\mu \epsilon \partial_\mu,$$

which has the component $K^u = \frac{1}{\sqrt{2}} (\Gamma^+ \epsilon)^T (\Gamma^+ \epsilon)$. The extra supersymmetries with $\Gamma^+ \epsilon \neq 0$ will give us the nonvanishing Killing component K^u .

2. A Killing vector with nonvanishing K^u component is not enough to ensure the extra supersymmetries. Since in our discussion the Killing spinor is independent of v , the corresponding Killing vector cannot have v dependence.

- **Results:** For the metric

$$ds^2 = 2A_0(u)dudv + B_{ij}(u)x^i x^j (du)^2 + A_i(u)(dx^i)^2$$

the possible time-dependent configurations with supernumerary supersymmetries are

- $B_0 = 0, A_0 = 1$, and A_i 's are exponential functions;
- $B_0 \neq 0, A_0 = 1$, A_i 's are exponential functions and B_{jk} take the form of

$$B_{jk} = e^{\gamma_j u} (e^{2uf} B_s e^{-2uf})_{jk}$$

where B_s is a constant symmetric matrix and f is a constant antisymmetric matrix.

However, after coordinate transformation, the above backgrounds are reduced to the well-studied homogeneous plane-waves. [Blau et.al.](#)

We conclude that our time-dependent supersymmetric configurations of the metric form (2.1) have no supernumerary supersymmetry, except the cases with constant A_0, A_i 's and the appropriately chosen B_{ij}, A_{ij} 's.

3 Matrix models

- Matrix models in the curved backgrounds is a subtle issue;
- Weak field limit following DLCQ prescription; [W. Taylor et.al.](#)

However, for the above time-dependent configurations, it could be far from flat near the big-bang singularity, and the validity of weak field limit is in doubt;

- Another way: matrix regularization of the membrane action in the curved background: order by order expansion in fermionic coordinates; [de Wit et.al.](#) People has applied this route to the flat spacetime and maximal plane-wave background and obtained the BFSS and BMN matrix model succesfully. [de Wit et.al., K. Dasgupta et.al.](#)
- In our case, we obtained the membrane action to all orders of fermionic coordinates. [H. Z. Chen and B. Chen](#)

The supermembrane action is:

$$S[Z(\xi)] = \int d^3\xi [-\sqrt{-g(Z(\xi))} - \frac{1}{6}\epsilon^{abc}\Pi_a^A\Pi_b^B\Pi_c^C B_{CBA}(Z(\xi))],$$

where $Z^A(\xi) = (x^\mu(\xi), \theta(\xi))$ is the curved superspace coordinates, $g_{ab} = \Pi_a^\mu\Pi_b^\nu g_{\mu\nu} = \Pi_a^r\Pi_b^s\eta_{rs}$ is the induced metric, $\eta_{rs} = \text{diag}(-1, 1, \dots, 1)$ is the 11-d Lorentz metric, and $\xi^a = (\xi^0, \xi^1, \xi^2) = (\tau, \xi^\alpha)$, $\alpha = 1, 2$ represent the coordinates on the world volume. Here Π_a^A are the supervielbein pullback, B_{ABC} are the super three-potential.

In our case, the gravitino is zero, so the supervielbein pull-back is:

$$\begin{aligned}\Pi_a^r &= \partial_a Z^A E_A^r \\ &= \partial_a x^\mu \left(e_\mu^r - \frac{1}{4} \bar{\theta} \Gamma^{rst} \theta \omega_{\mu st} + \bar{\theta} \Gamma^r \Omega_\mu \theta \right) + \bar{\theta} \Gamma^r \partial_a \theta + \mathcal{O}(\theta^3),\end{aligned}$$

where $\omega_{\mu st}$ is the spin connection, and

$$\Omega_\mu = \frac{1}{288} F_{\nu\rho\sigma\lambda} (\Gamma_\mu^{\nu\rho\sigma\lambda} + 8\Gamma^{\nu\rho\sigma} \delta_\mu^\lambda).$$

The super three-potential pullback is:

$$\begin{aligned}
& - \frac{1}{6} \Pi_a^A \Pi_b^B \Pi_c^C B_{CBA} \\
& = \frac{1}{6} \epsilon^{abc} \partial_a x^\mu \partial_b x^\nu \partial_c x^\rho \left[C_{\mu\nu\rho} + \frac{3}{4} \bar{\theta} \Gamma_{rs} \Gamma_{\mu\nu} \theta \omega_\rho^{rs} - 3 \bar{\theta} \Gamma_{\mu\nu} \Omega_\rho \theta \right] \\
& - \epsilon^{abc} \bar{\theta} \Gamma_{\mu\nu} \partial_c \theta \left[\frac{1}{2} \partial_a x^\mu (\partial_b x^\nu + \bar{\theta} \Gamma^\nu \partial_b \theta) + \frac{1}{6} \bar{\theta} \Gamma^\mu \partial_a \theta \bar{\theta} \Gamma^\nu \partial_b \theta \right] + \mathcal{O}(\theta^3),
\end{aligned}$$

where $C_{\mu\nu\rho}$ is the three-form potential.

In our configurations, the supervielbein and super three-potential are exact to all orders of θ since the higher order terms vanish. Why?

- The supervielbein pullback $\Pi_a^r = \partial_a Z^A E_A^r$ is linear in $\partial_a X^\mu$, while E_A^r is constituted with other quantities. It can be seen from their explicit form that these other quantities, θ , Γ^r , Ricci tensor, Ω_μ , and field strength et.al. have no lower curved spacetime index v , and hence no upper curved spacetime index u .

- The nonvanishing bilinear fermionic terms $\bar{\theta}\Gamma^{rst\dots}\theta$ always have one and only one Γ^- and no Γ^+ due to the gauge condition $\Gamma^+\theta = 0$. The upper tangent space index $r = -$ require an upper curved spacetime index $\mu = u$ coming from other geometrical quantities because the only nonzero vielbein with a lower tangent index $r = -$ is e_{-u} . Such an index cannot be cancelled by the above mentioned quantities except $\partial_a X^u$.

- the super three-potential pullback term can only have bilinear θ terms. This is due to the antisymmetric nature of ϵ^{abc} and the fact that bilinear θ term in Π_a^A must be proportional to $\partial_a X^u$.

To simplify the action, we go to light-cone gauge:

$$x^u = u = \tau.$$

And because of the κ -symmetry of the action, we can also impose an additional gauge

$$\Gamma^+ \theta = 0.$$

There still exist the residue gauge D.O.F. We should fix them and write our action in pure physical D.O.F.

Let us focus on the following metric and field strength

$$\begin{aligned}
 ds^2 &= 2e^{r_0 u} du dv + \sum_i c_i e^{r_i u} (x^i)^2 (du)^2 + \sum_i e^{r_i u} (dx^i)^2 \\
 &+ \sum_{ij} A_{ij}^0 e^{(r_i+r_j)u/2} x^j dx^i du,
 \end{aligned}$$

where

$$A_{ij}^0 = -A_{ji}^0 = \text{const},$$

and r_0, r_i are all constants, too. We also have a four-form field strength

$$F_{u123} = e^{(r_1+r_2+r_3)u/2} f^0, \quad f^0 = \text{const}.$$

The Lagrangian of the membrane action in this background is

$$\begin{aligned}
\mathcal{L} = & \sum_i \frac{P^u}{2} e^{(r_i - r_0)\tau} (D_\tau x^i)^2 + \frac{P^u}{2} \sum_{ij} A_{ij}^0 e^{(\frac{r_i + r_j}{2} - r_0)\tau} x^j D_\tau x^i \\
& + \frac{P^u}{2} \sum_i c_i e^{(r_i - r_0)\tau} (x^i)^2 - \frac{e^{r_0\tau}}{4P^u} \sum_{ij} e^{(r_i + r_j)\tau} \{x^i, x^j\}^2 \\
& - \frac{1}{2} \sum_{i,j=1,2,k=3} \{x^i, x^j\} x^k \epsilon_{ijk} f^0 e^{(r_i + r_j + r_k)\tau/2} \\
& + iP^u e^{-r_0\tau/2} \psi^T D_\tau \psi - \frac{i}{6} P^u e^{-r_0\tau/2} f^0 \psi^T \gamma^{123} \psi \\
& - \frac{i}{8} P^u \sum_{IJ} e^{-r_0\tau/2} A_{IJ}^0 \psi^T \gamma^{IJ} \psi - i \sum_{I,i} \psi^T \gamma^I \{x^i, \psi\} e^{(r_0 + r_i)\tau/2} \delta_i^I.
\end{aligned}$$

The usual matrix regularization:

$$x^i \rightarrow X_{N \times N}^i,$$

$$\psi \rightarrow \psi_{N \times N},$$

$$P^u \int d^2\sigma \rightarrow \frac{1}{R} \text{Tr},$$

$$\{, \} \rightarrow -i[,],$$

Then the matrix model action reads

$$\begin{aligned}
S = & \int d\tau \text{Tr} \left(\sum_i \frac{1}{2R} e^{(r_i - r_0)\tau} (D_\tau X^i)^2 + \frac{1}{2R} \sum_{ij} A_{ij}^0 e^{(\frac{r_i + r_j}{2} - r_0)\tau} X^j D_\tau X^i \right. \\
& + \frac{1}{2R} \sum_i c_i e^{(r_i - r_0)\tau} (X^i)^2 + \frac{R}{4} e^{r_0\tau} \sum_{ij} e^{(r_i + r_j)\tau} [X^i, X^j]^2 \\
& + \frac{i}{2} \sum_{i,j=1,2,k=3} [X^i, X^j] X^k \epsilon_{ijk} f^0 e^{(r_i + r_j + r_k)\tau/2} \\
& + \frac{i}{R} e^{-r_0\tau/2} \psi^T D_\tau \psi - \frac{i}{6R} e^{-r_0\tau/2} f^0 \psi^T \gamma^{123} \psi \\
& \left. - \frac{i}{8R} \sum_{IJ} e^{-r_0\tau/2} A_{IJ}^0 \psi^T \gamma^{IJ} \psi - \sum_{I,i} \psi^T \gamma^I [X^i, \psi] e^{(r_0 + r_i)\tau/2} \delta_i^I \right).
\end{aligned}$$

- **Fuzzy sphere solutions:** consider the matrix model in the sector:

$$X^4 = X^5 = \dots = X^9 = 0, \psi = 0.$$

To simplify the problem, let

$$r_1 = r_2 = r_3 = r, c_1 = c_2 = c_3 = c.$$

We want to find solution of the form:

$$X^a(\tau) = S(\tau)J^a, a = 1, 2, 3,$$

where J^a is N dimensional representation of $SU(2)$.

Use

$$\text{Tr} \sum_a (J^a)^2 = \frac{N(N-1)}{4},$$

and

$$[J^a, J^b] = i\epsilon^{abc} J^c.$$

We finally get

$$\frac{d^2 S}{d\tau^2} + (r - r_0) \frac{dS}{d\tau} + 2R^2 e^{(2r_0+r)\tau} S^3 + Rf^0 e^{(r_0+r/2)\tau} S^2 - cS = 0.$$

4 IIB time-dependent plane-wave with linear null dilaton

- Linear null dilaton is different from the usual linear (spatial) dilaton;
- Plane-fronted waves are exact solutions of the string theory; [Horowitz and Stief](#)
- Plane-wave/SYM correspondence. [Metsaev et.al.](#), [BMN](#)

Let's consider the IIB GS superstrings in the following backgrounds:

$$ds^2 = -2dx^+ dx^- - \lambda(x^+) x_I^2 dx^+ dx^+ + dx^I dx^I ,$$

$$\phi = \phi(x^+) , \quad (F_5)_{+1234} = (F_5)_{+5678} = 2f .$$

The world-sheet conformal invariance requires

$$\lambda = -\frac{1}{4}\phi'' + f^2 e^{2\phi} .$$

If we restrict the dilaton to be linear in the light-cone time coordinate x^+ , i.e. $\phi = -cx^+$ with c being a constant, then we have

$$\lambda = f^2 e^{-2cx^+}.$$

In general, f could be an arbitrary function of x^+ and there is a large class of the models. These models could be studied in the lightcone gauge.

- Two special cases:

1. One is that λ is a constant. This happens when $f = f_0 e^{-\phi}$ with f_0 being constant so that

$$\lambda = f_0^2,$$

and the metric in the string frame reduces to the form of the maximally supersymmetric plane wave;

[Blau et.al.](#), [Mestaeu et.al.](#), [BMN](#)

2. The other special case is when the self-dual RR field strength $f = f_0$ is constant so that

$$\lambda = f^2 e^{-2cx^+}.$$

- **Symmetry algebra:**

1. A null Killing vector T , corresponding to the translation in the x^- ;
2. Two $SO(4)$ rotations $+ Z_2$;
3. Translation along x^I plus the shift along x^- ;

The the continuous symmetry algebra is $[so(4) \oplus so(4)] \oplus_s h(8)$.

- **Implications:** light-cone momentum conserved; no particle or string creation; no supernumerary supersymmetries; Hamiltonian not conserved.

- Geodesics incomplete

Let us focus on $x^I = 0, x^- = \text{constant}$, and consider the geodesic equation for x^+

$$\frac{d^2 x^+}{d\sigma^2} + \frac{c}{2} \left(\frac{dx^+}{d\sigma} \right)^2 = 0 ,$$

which gives

$$\sigma = e^{\frac{c}{2}x^+}$$

up to an affine transformation. Therefore the singularity $x^+ \rightarrow -\infty$ corresponds to $\sigma = 0$ and it has finite affine distance to all points in the interior.

- **Supersymmetries**

The gravitino and dilatino variations should vanish for independent Killing spinors, i.e.

$$\delta_\epsilon \lambda^A \equiv (\tilde{\mathcal{D}})^A_B \epsilon^B = 0, \quad \delta_\epsilon \psi_\mu^A \equiv (\hat{\mathcal{D}}_\mu)^A_B \epsilon^B = 0,$$

where $\mu = +, -, 1, \dots, 8$ and $A = 1, 2$ and

$$\begin{aligned} (\tilde{\mathcal{D}})^A_B &= \frac{1}{2} \delta_B^A \Gamma^\mu \partial_\mu \phi, \\ (\hat{\mathcal{D}}_\mu)^A_B &= \delta_B^A \partial_\mu + (\Omega_\mu)^A_B, \end{aligned}$$

with

$$(\Omega_\mu)_B^A = \frac{1}{4} \omega_\mu^{\hat{\nu}\hat{\rho}} \Gamma_{\hat{\nu}\hat{\rho}} \delta_B^A + \frac{i e^\phi}{8 \cdot 5!} \Gamma^{\kappa\nu\rho\sigma\delta} F_{\kappa\nu\rho\sigma\delta} \Gamma_\mu (\sigma_2)_B^A,$$

where σ_2 is the Pauli matrix, $\omega_\mu^{\hat{\nu}\hat{\rho}}$ is the spin connection and the hatted indices are used for the tangent space.

It's straightforward to check our background keep one-half of the original supersymmetries.

- **Bosonic sector**

The bosonic part of the GS action is

$$S_B = -\frac{1}{4\pi\alpha'} \int d^2\sigma \sqrt{-g} g^{ab} G_{\mu\nu} \partial_a x^\mu \partial_b x^\nu$$

In the light-cone gauge the bosonic action takes the form

$$S_B = \frac{1}{4\pi\alpha'} \int d\tau \int_0^{2\pi} d\sigma (\partial_\tau x^I \partial_\tau x^I - \partial_\sigma x^I \partial_\sigma x^I - \tilde{f}^2 e^{-2\tau} x_I^2) ,$$

where for simplicity we have defined $\tilde{f} \equiv \alpha' p^+ f$ which is dimensionless.

The equations of motions are

$$(\partial_\tau^2 - \partial_\sigma^2 + \tilde{f}^2 e^{-2\tau}) x^I = 0.$$

Expanding in Fourier modes in σ , we get an infinite collection of oscillators with time-dependent frequencies.

The bosonic light-cone Hamiltonian is

$$H_B = H_{B0}(\tau) + \frac{1}{\alpha' p^+} \sum_{n=1}^{\infty} \omega_n(\tau) \left[A_{-n}^I(\tau) A_n^I(\tau) + \tilde{A}_{-n}^I(\tau) \tilde{A}_n^I(\tau) + 1 \right],$$

where

$$\omega_n = \sqrt{n^2 + \tilde{f}^2 e^{-2\tau}}. \quad (4.1)$$

The bosonic Hamiltonian reminisces the 2-d field theory of free scalars with time-dependent masses.

- **Fermionic sector**

The fermionic action in the light-cone gauge can be written as

$$S_F = \frac{ip^+}{\sqrt{2}\pi} \int d\tau \int_0^{2\pi} d\sigma (\theta^{1T} \partial_\tau \theta^1 + \theta^{2T} \partial_\tau \theta^2 + \theta^{1T} \partial_\sigma \theta^1 - \theta^{2T} \partial_\sigma \theta^2 + 2\tilde{f} e^{-\tau} \theta^{1T} \Pi \theta^2).$$

Similar to the bosonic case, when $f = f_0 e^{-\phi}$, the action reduces to the one in the usual plane-wave metric.

The equations of motion of the fermionic sector are

$$\begin{aligned}(\partial_\tau + \partial_\sigma) \theta^1 + \tilde{f} e^{-\tau} \Pi \theta^2 &= 0, \\(\partial_\tau - \partial_\sigma) \theta^2 - \tilde{f} e^{-\tau} \Pi \theta^1 &= 0.\end{aligned}$$

The fermionic part of the Hamiltonian can be diagonalized as

$$H_F = H_{F0}(\tau) + \frac{1}{\alpha' p^+} \sum_{n=1}^{\infty} \omega_n(\tau) \left[B_n^\dagger(\tau) B_n(\tau) + \tilde{B}_n^\dagger(\tau) \tilde{B}_n(\tau) - 1 \right].$$

Just like the bosonic case, it also looks like the Hamiltonian of a free 2-d field theory with time-dependent mass.

- **Quantum string mode creation**

Generically, in a time-dependent background, one may expect the particle or string creation occurs from our knowledge of the quantum field theory in curved space-time. However, in a plane-fronted background, due to the existence of null Killing vector, this would not happen. **Gibbons 1975**

Nevertheless, there does exist the string mode creation. **Horowitz 1990**

In our case, the total number of created oscillator modes is

$$\bar{N}_T(\tau) = \sum_{n=1}^{\infty} [\bar{N}_n^B + \bar{N}_n^F].$$

1. As $\tau \rightarrow +\infty$, $\bar{N}_T(\tau) \sim 0$;
2. As $\tau \rightarrow -\infty$, the total number of created oscillator modes is

$$\begin{aligned} \bar{N}_T(\tau) &= 4d \sum_{n=1}^{\infty} \frac{1}{e^{2n\pi} - e^{-2n\pi}} \\ &\sim 0.06 . \end{aligned}$$

Effectively, the problem could be restated as a quantum mechanical problem. A Bosonic mode $T_n(\tau)$ should satisfy the equation

$$\partial_\tau^2 T_n + (n^2 + \tilde{f}^2 e^{-2\tau}) T_n = 0.$$

Replacing τ by x and T_n by ψ , the above equation takes a form of one-dimensional Schrodinger equation for a particle with energy n^2 in a potential $-\tilde{f}^2 e^{-2\tau}$. The problem of calculating the number of the creating modes reduces to the problem of calculating the reflective amplitude in this one-dimensional system.

- **Remarks**

1. The symmetric spectrum of the bosonic and fermionic excitations does not come from the spacetime supersymmetries. It is from the special choice of the metric. Consider the following background

$$ds^2 = -2dx^+ dx^- - \sum_I \lambda_I(x^+) x_I^2 dx^+ dx^+ + dx^I dx^I ,$$
$$\phi = \phi(x^+) , \quad (F_5)_{+1234} = (F_5)_{+5678} = 2f .$$

Here λ_I could be different from each other. The transverse bosons in the bosonic action have different masses proportional to λ_I .

2. Matrix description near the big-bang singularity? There exists the cosmological singularity at $x^+ = -\infty$, where the string coupling is divergent. (Matrix model in IIB ?)
3. Membrane creation? [S. Das et.al](#)

- Other directions?

Holographic description of the time-dependent backgrounds

Chu and Ho, S.Das et.al., F.L. Lin et.al.

(S)YM with time-dependent coupling, what does it mean?
nontrivial RG flow? Exotic QFT!

Why the correspondence could be true? Short of evidence!

Stability? 1-loop issue.